







Missione Rosetta: come raggiungere ed esplorare una cometa

Politecnico di Milano

Dipartimento di Scienze e Tecnologie Aerospaziali

Prof. Franco Bernelli

The Rosetta mission A «staggeringly ambitious plan»

- It's the first space mission ever launched to:
 - chase
 - orbit around
 - land on
 - sample

a comet



Courtesy of ESA

- perform scientific observation of:
 - o comet's nucleus
 - o **coma**





Why targetting a comet?

Because...

- Comets are made of *primordial materials*
- Comets remain *inert* for a very long time → small changes expected in their original formation
- Comets transport materials (*organic*? *life building blocks*?) from one side to the other of the Solar System



The Rosetta space mission main objective

(citation from sci.esa.int/rosetta)

To study the origin of comets, the relationship between cometary and interstellar material, and its implications with regard to the origin of the Solar System

translates into...

- "Global characterisation of the nucleus, determination of dynamic properties, surface morphology and composition
- determination of the chemical, mineralogical and isotopic
 compositions of volatiles and refractories in a cometary nucleus"



- "Determination of the physical properties and interrelation of volatiles and refractories in a cometary nucleus
- Study of the development of cometary activity and the processes in the surface layer of the nucleus and the inner coma (dust/gas interaction)
- 5. Global **characterisation of asteroids**, including determination of dynamic properties, surface morphology and composition."

(citation from sci.esa.int/rosetta)

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Rosetta's final target Comet Churymov-Gerasimenko (C-G) / 67P



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- \checkmark an **orbiter** \rightarrow
- Rosetta, with 11 scientific experiments
- ✓ a lander \rightarrow
- *Philae*, with 10 scientific instruments for in situ analysis of the comet surface



The orbiter

- ✓ The case to protect \rightarrow
- \checkmark To talk \rightarrow
- ✓ The energy source →

- body size: 2.8 x 2.1 x 2.0 m
- high gain antenna diameter: 2.2 m
 - photovoltaic Si cells: 2wings, 32m² wide each, still providing 400W at ~760<u>000000 km from the Sun!!</u>
- ✓ Launch mass ~ 3000 kg:
 - Propellant: 1670 kg
 - Science payload: 165 kg
 - Lander: 100 kg
- ✓ Max power demand 900W



Missions to comets

- ISEE-3/ICE
- Vega 1 and 2
- Sakigake
- Suisei
- Giotto

Stardust

- Stardust
- Deep Impact



➡ Halley (1986)

> Wild 2 (2004)

Tempel 1 (2005)







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- Rosetta mission was:
 - conceived in the late 1970s
 - approved in 1993 as a Cornerstone Mission
 - Launched in 2004 by Ariane 5 rocket, form Kourou French Guiana



Courtesy of ESA

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Rosetta Mission: history

First conceived in the late '70s to explore comet 46P/Wirtanen



Champollion (France, USA)



RoLand (Germany, Italy)



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Rosetta Mission: history

Approved by ESA in the early '90s to explore comet 46P/Wirtanen





Philae (Germany, France, Italy)

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Rosetta Mission: launch

Planned launch: March 2003

Ariane 5 failure: 11 December 2002







Planned launch: March 2003

46P / Wirtanen



Actual launch: 02/03/2004

67P/Churyumov-Gerasimenko



The consortium for the design and construction phases



15

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Solar Array Assembly



Courtesy of Selex ES

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Philae Lander: scientific instruments





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Drill prototype test, Etna, 2002





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Rosetta mission timeline

Event	Date
Launch	March 2004
1 st Earth Gravity Assist (GA)	March 2005
Mars GA	February 2007
2 nd Earth GA	November 2007
Flyby of Asteroid Steins	September 2008
3 rd Earth GA	November 2009
Flyby of Asteroid Lutetia	July 2010
Hybernation	July 2011
Wake-up	January 2014
Comet rendezvous	May 2014
Philae Landing	November 2014
Comet Closest Approach to the Sun	August 2015
Nominal Mission End	December 2015

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OSIRIS WAC Ultraviolet image 24 February 2007

Clouds in Mars atmosphere

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CIVA Image 25 February 2007 About 1000 km distance

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NAVCAM image 13 November 2007 About 6250 km from surface

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A diamond in space OSIRIS image – 5 September 2008 – 800 km distance







OSIRIS image November 2009 About 350000 km distance

Courtesy of ESA



VIRTIS images 13 November 2009 About 230000 km distance



Courtesy of ESA

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OSIRIS images – 10 July 2010 – 51000-81000 km distance





OSIRIS images 10 July 2010 Closest approach – 3<u>160km</u>



Courtesy of ESA

Rosetta close encounter with asteroid Lutetia



Courtesy of ESA

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Lamy et al., 2007



Figure 5. The prograde A1 (top row) and the retrograde B2 (bottom row) solutions for the threedimensional shape of the nucleus of comet 67P/Churyumov–Gerasimenko reconstructed from the inversion of the 2003 HST and 2005 NTT light curves. For each solution, three views of the reconstructed 3-D shape model are displayed at three different rotational phase angles: 350° (left-panel), 80° (mid-panel), and pole-on view of the 80° model (right-panel).

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February 2

2014



July 4



August 3

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credits: ESA/OSIRIS

32



From on acquired scientific data, preliminary infos on the target have been obtained

Mass:	10 ¹³ kg
Volume:	25 km ³
Density:	0.4 g/cm ³
Rotational period:	12.4 hours



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Courtesy of ESA





Courtesy of ESA









Image from CIVA -7 October from a distance of about 16 km from the comet

Courtesy of ESA





credits: ESA

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ESA/Rosetta/MPS for OSIRIS Team MPS/UPD/LAM/IAA/SSO/INTA/UPM/DASP/IDA

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Rosetta Mission Control Center (Darmstatd) ...





46

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November 14, 2014: SD2

- SD2 procedure:
- 1. Drill to the position 250 mm
- 2. Extract the sampling tube and rotate to perform a coring
- 3. Translate back to home position
- 4. Rotate Carousel to move the desired oven under the drill
- 5. Discharge the sample into the oven
- 6. Rotate Carousel to move the oven with its sample under **COSAC**
- 7. Rearm the drill bit before the next drilling and sampling activity



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Final sequence of operations







- No mechanical movements
 - ROMAP: Magnetometer
 - CONSERT: Sounding Experiment by Radiowave Transmission
 - MUPUS: temperature mapping

Final sequence of operations



- No mechanical movements
 - ROMAP: magnetometer
 - MUPUS: temperature mapping
 - SESAME-DIM: dust impacts monitor
 - COSAC and PTOLEMY sniffing





- Mechanical movements
 - MUPUS: hammering mode to measure compressive strength
 - APXS: Alpha proton X-ray spectrometer to measure composition
 - SESAME-CASSE: sounding experiment (seismograph and sonar)

Final sequence of operations



- Mechanical movements
 - SD2: drilling and sample collection
 - COSAC: sample analysis
 - Lander rotation to increase power production
 - PTOLEMY: sniffing of the Carbosphere oven
 - ROLIS: image of the soil after lander rotation
 - CONCERT: sounding experiment by radiowave transmission

Communications with Philae











November 14, 2014

















62

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